

Today 13.3, 13.4

L36



Today 13.3, 13.4

L36

Gravitational  
potential  
energy



Today 13.3, 13.4

L36

Gravitational  
potential  
energy

Motion  
of satellites

Today 13.3, 13.4

L36

AI session

Today 13.3, 13.4

L36

AI session

\* 5pm to 6pm

Today 13.3, 13.4

L36

## TI session

\* 5pm to 6pm

\* §13.1 Newton's law of gravitation

Today 13.3, 13.4

L36

## TI session

\* 5pm to 6pm

\* §13.1 Newton's law of gravitation

\* §13.2 Weight

Today 13.3, 13.4

Wednesday 13.5, 13.8

L36



Today 13.3, 13.4

Wednesday 13.5, 13.8

L36

Kepler's  
Laws

Today 13.3, 13.4

Wednesday 13.5, 13.8

Kepler's  
laws

Black  
holes

Today 13.3, 13.4

Wednesday 13.5, 13.8

Friday Holiday

L36

L36

Today 13.3, 13.4

Wednesday 13.5, 13.8

Friday Holiday

Monday Nov. 30<sup>th</sup> Exam # 4



Today 13.3, 13.4

Wednesday 13.5, 13.8

Friday Holiday

Monday Nov. 30<sup>th</sup> Exam # 4

Wednesday Dec. 2<sup>nd</sup> Day of Reckoning

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Monday Nov. 30<sup>th</sup> Exam # 4

Wednesday Dec. 2<sup>nd</sup> Day of Reckoning

Friday Dec. 4<sup>th</sup> Final exam



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In these notes I mix some notation 😞

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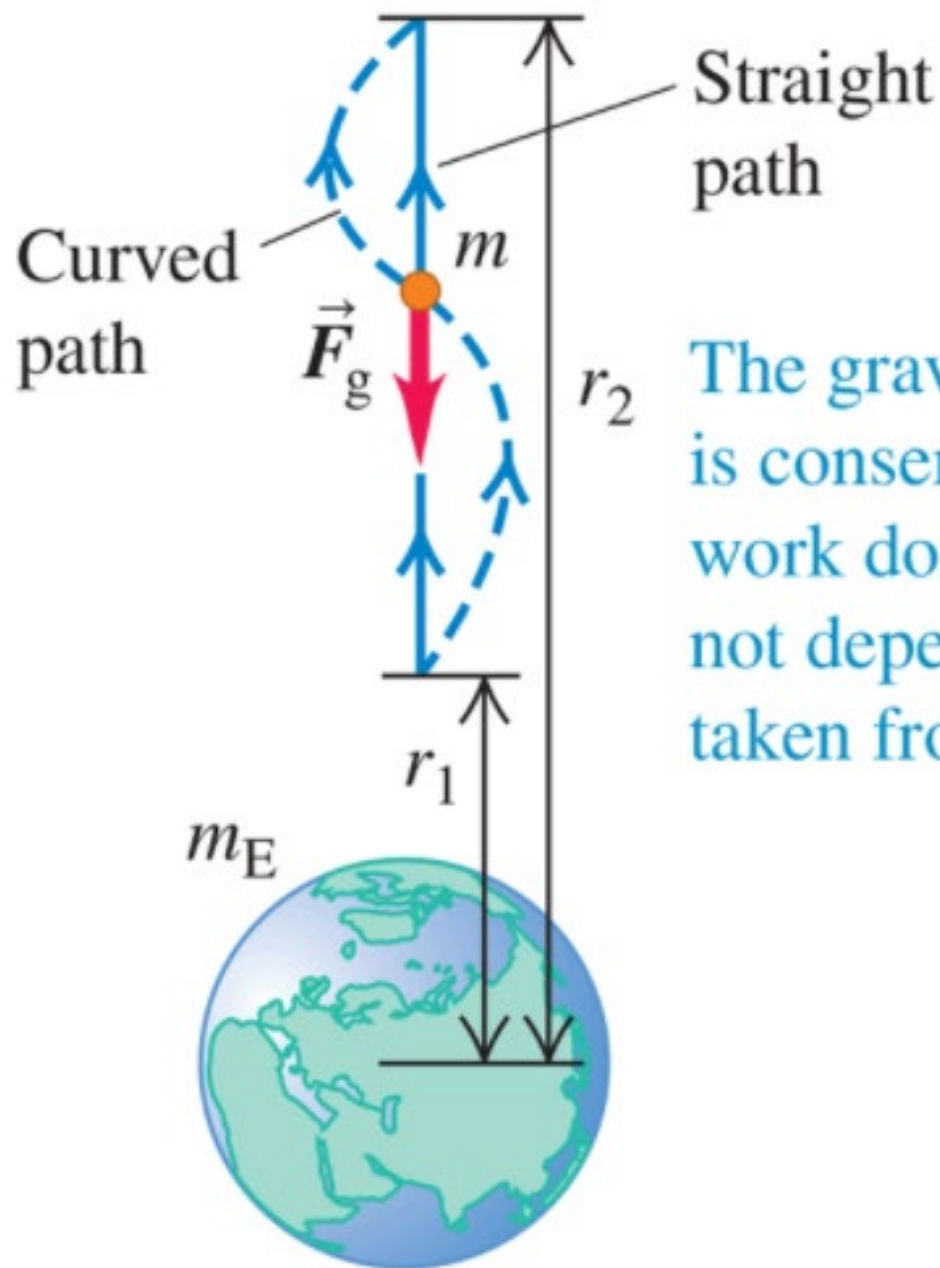
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$U \rightarrow$  Potential energy

& Sometimes I use

$V \rightarrow$  Potential energy



The gravitational force is conservative: The work done by  $\vec{F}_g$  does not depend on the path taken from  $r_1$  to  $r_2$ .

$$W_{\text{grav}} = \int_{r_1}^{r_2} F_r dr$$

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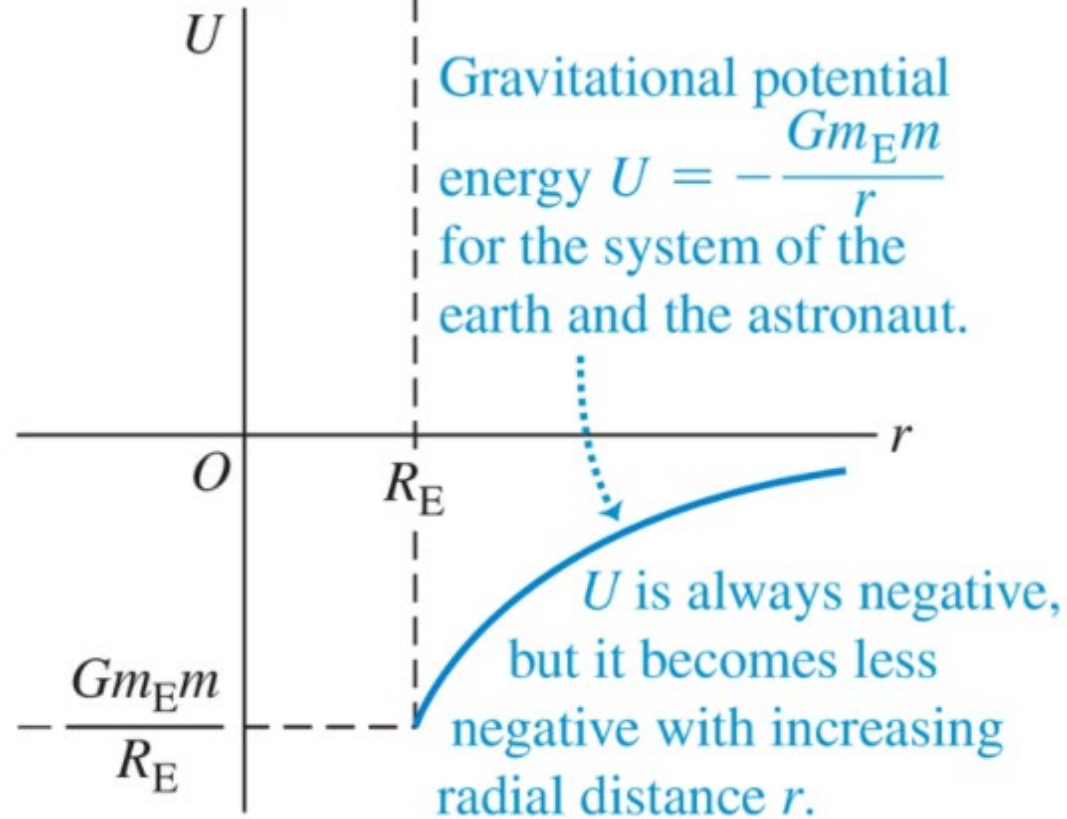
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Earth, mass  $m_E$



Astronaut, mass  $m$



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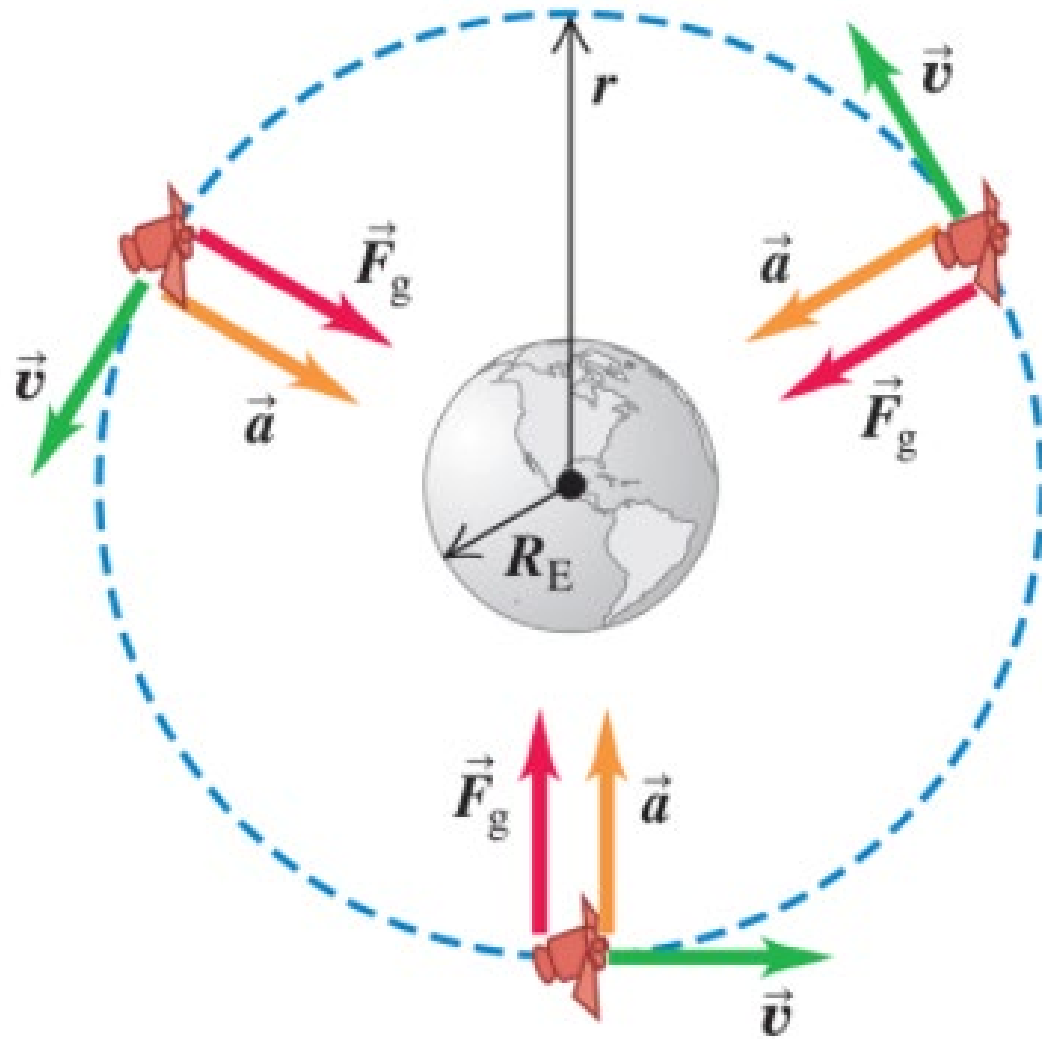
$$v_1 = \sqrt{2gR_E} = [2 * 9.8 * 6.37]^{1/2} * 10^3 \text{ m/s}$$

$$\Rightarrow v_1 = 11 * 10^3 \text{ m/s}$$

§13.4

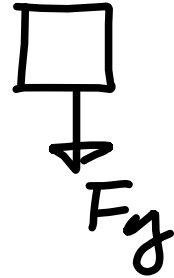
Motion of Satellites





The satellite is in a circular orbit: Its acceleration  $\vec{a}$  is always perpendicular to its velocity  $\vec{v}$ , so its speed  $v$  is constant.

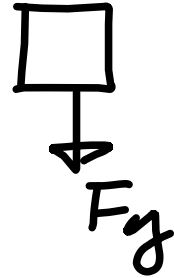
$$\Sigma F_{rad} = M a_{rad}$$



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⇒

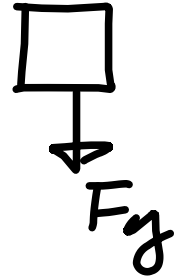
$$F_g = M \frac{v^2}{r}$$



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⇒

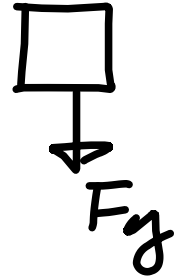
$$F_g = M \frac{v^2}{r} \Rightarrow \frac{GM_E M}{r^2} = M \frac{v^2}{r}$$



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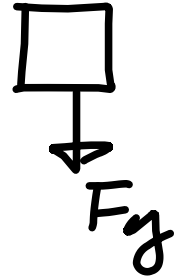
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⇒

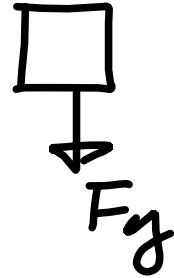
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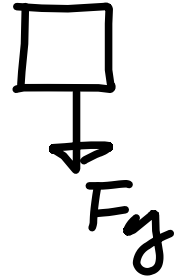
$$\Rightarrow v^2 = \frac{GM_E}{r}$$



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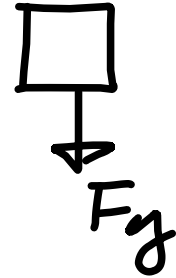
$$\Rightarrow v^2 = \frac{GM_E}{r} \Rightarrow v = \sqrt{\frac{GM_E}{h}}$$



$$\Sigma F_{rad} = M a_{rad}$$

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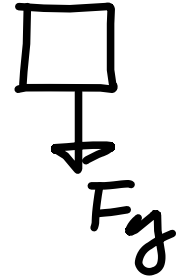


Since  $v = \frac{2\pi r}{T}$

$$\Sigma F_{rad} = M a_{rad}$$

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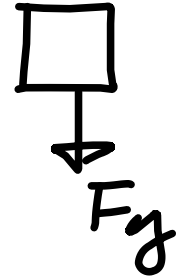
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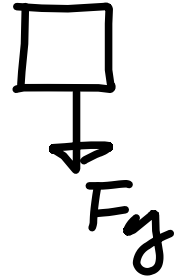
Since  $v = \frac{2\pi r}{T}$ , then

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For  
circular  
orbit

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$$T = \left( \frac{2 \times 3.14 \times 6710 \times 10^3}{7.7 \times 10^3} \right) \frac{\text{m}}{\text{s}}$$



You wish to put a 1000 kg satellite into a circular orbit 340 km above the earth's surface. (a) What speed, period, and radial acceleration will it have? (b) How much work must be done to the satellite to put it in orbit? (c) How much additional work would have to be done to make the satellite escape the earth? The earth's radius and mass are given in Example 13.5

(Section 13.3).  $M = 1000 \text{ kg}$ ,  $R = R_E + 340 \text{ km}$ .

Find  $v$ ,  $T$ ,  $a_{\text{rad}}$ :  $\Sigma F = ma \Rightarrow F_g = M \frac{v^2}{R}$

$$\Rightarrow v^2 = \frac{R F_g}{M} \Rightarrow v^2 = \left( \frac{R}{M} \right) \left( \frac{G M_E M}{R^2} \right) \Rightarrow$$

$$v = \sqrt{G M_E / R} = \left[ \frac{(6.67 \times 10^{-11})(5.97 \times 10^{24})}{(6370 + 340) \times 10^3} \right]^{1/2} \frac{\text{m}}{\text{s}}$$

$$= \left[ \frac{(6.67)(5.97)}{6710} \right]^{1/2} \times 10^{(24-11-3)/2} \frac{\text{m}}{\text{s}} = 0.077 \times 10^5 \frac{\text{m}}{\text{s}}$$

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$a_{\text{rad}} = \frac{v^2}{R}$



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$$a_{rad} = \frac{v^2}{R} \Rightarrow$$

$$a_{rad} = \frac{v^2}{R} \Rightarrow a_{rad} = \frac{[7.7 * 10^3]^2}{6710 * 10^3} \left( \frac{m}{s^2} \right)$$

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(Section 13.3).  $m = 1000 \text{ kg}$ ,  $R = R_E + 340 \text{ km}$

$$W = E_2 - E_1$$

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$$\Rightarrow W = 9.8 * 1000 * 6340^2 * 10^6 \left( \frac{1}{6340} - \frac{1}{2 * 6370} \right) \frac{\text{J}}{10^3}$$

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$$F_g = M V_2^2 / R \Rightarrow M V_2^2 = R F_g = \frac{GM_E M}{R} \Rightarrow K_2 = \frac{GM_E M}{2R}$$

$$\text{Now } W = K_2 + V_2 - V_1 = (GM_E M) \left[ \frac{1}{2R} + \left(-\frac{1}{R}\right) - \left(-\frac{1}{R_E}\right) \right]$$

$$\Rightarrow W = (GM_E M) \left[ \frac{1}{R_E} - \frac{1}{2R} \right] = (g M R_E^2) \left( \frac{1}{R_E} - \frac{1}{2R} \right)$$

$$\Rightarrow W = 9.8 * 1000 * 6340^2 * 10^6 \left( \frac{1}{6340} - \frac{1}{2 * 6370} \right) \frac{\text{J}}{10^3}$$

$$\Rightarrow W = 3.28 * 10^7 * 10^3 \text{ J} \Rightarrow W = 3.28 * 10^{10} \text{ J}$$



You wish to put a 1000 kg satellite into a circular orbit 340 km above the earth's surface. (a) What speed, period, and radial acceleration will it have? (b) How much work must be done to the satellite to put it in orbit? (c) How much additional work would have to be done to make the satellite escape the earth? The earth's radius and mass are given in [Example 13.5](#)

([Section 13.3](#)).  $m = 1000 \text{ kg}$ ,  $R = R_E + 340 \text{ km}$

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$$W = mg \frac{R_E^2}{2R} = \frac{1000 * 9.8 * 6370^2 * 10^6 \text{ J}}{2 * 6710 * 10^3} = 2.96 * 10^6 \text{ J}$$



