

Today 13.1, 13.2

235



Today 13.1, 13.2

235

Newton's
law of
gravitation



Today 13.1, 13.2

235

Newton's
law of
gravitation

Weight



Today 13.1, 13.2

Monday 13.3, 13.4

L35



Today 13.1, 13.2

235

Monday 13.3, 13.4

Gravitational
potential
energy



Today 13.1, 13.2

235

Monday 13.3, 13.4

Gravitational
potential
energy

Motion of
satellites



Today 13.1, 13.2

Monday 13.3, 13.4

Wednesday 13.5, 13.8

235



Today 13.1, 13.2

Monday 13.3, 13.4

Wednesday 13.5, 13.8

Kepler's
Laws

Today 13.1, 13.2

Monday 13.3, 13.4

Wednesday 13.5, 13.8

Kepler's
laws

Black
holes

Today 13.1, 13.2

Monday 13.3, 13.4

Wednesday 13.5, 13.8

Friday Nov. 27th Holiday



Today 13.1, 13.2

Monday 13.3, 13.4

Wednesday 13.5, 13.8

Friday Nov. 27th Holiday

Monday Nov. 30th Exam #4



Today 13.1, 13.2

Monday 13.3, 13.4

Wednesday 13.5, 13.8

Friday Nov. 27th Holiday

Monday Nov. 30th Exam #4

Wednesday Dec. 2nd Day of Reckoning

Today 13.1, 13.2

Monday 13.3, 13.4

Wednesday 13.5, 13.8

Friday Nov. 27th Holiday

Monday Nov. 30th Exam #4

Wednesday Dec. 2nd Day of Reckoning

Friday Dec 4th Final exam

Newton's Law of gravitation

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$$F_g = \frac{Gm_1m_2}{r^2}$$

Newton's Law of gravitation

Newton's law of gravitation:

Magnitude of attractive
gravitational force between
any two particles

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Gravitational constant (same for any two particles)

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Masses of particles

Newton's Law of gravitation

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Magnitude of attractive
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Gravitational constant (same for any two particles)

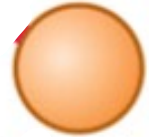
F_g

$$= \frac{Gm_1m_2}{r^2}$$

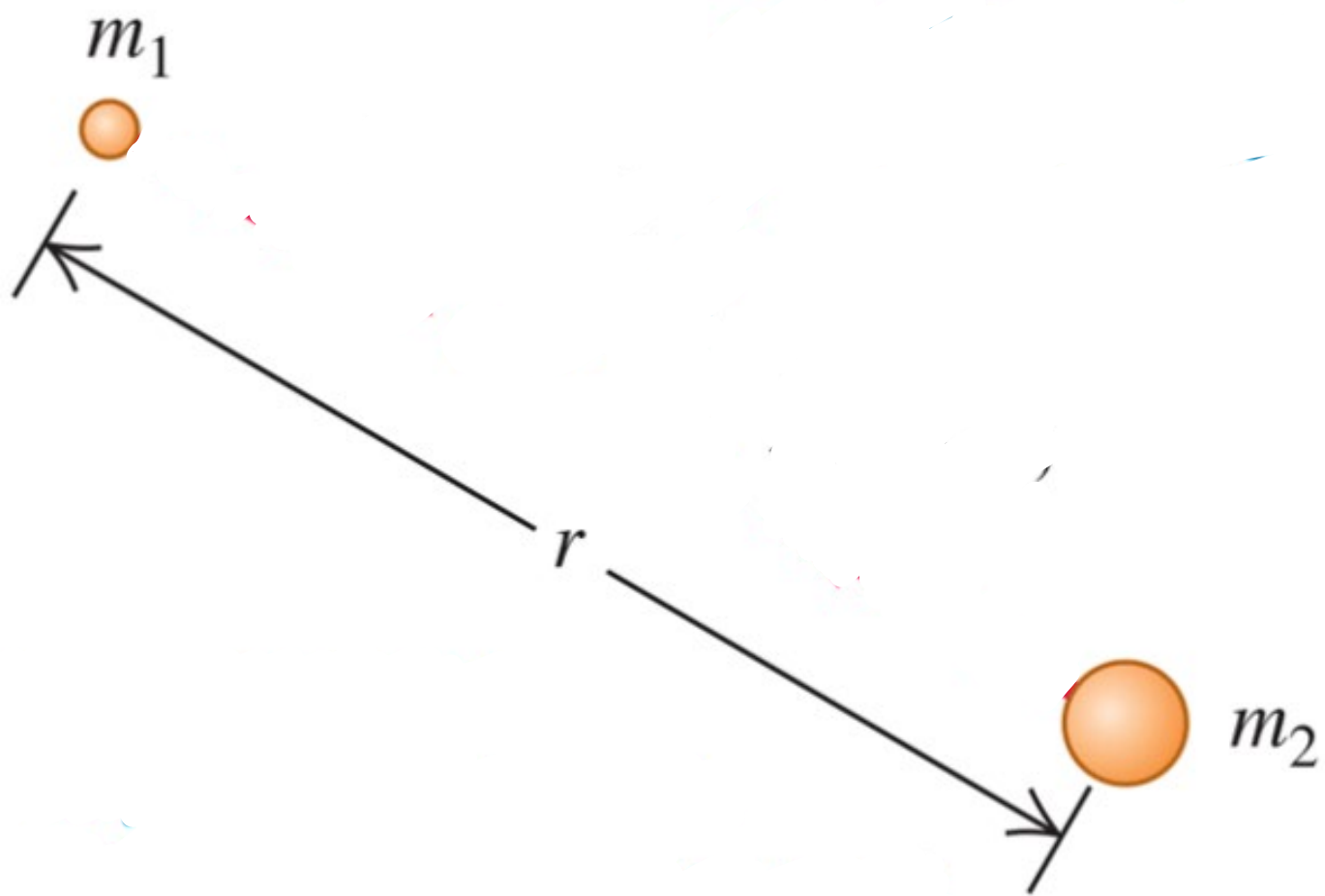
Masses of particles

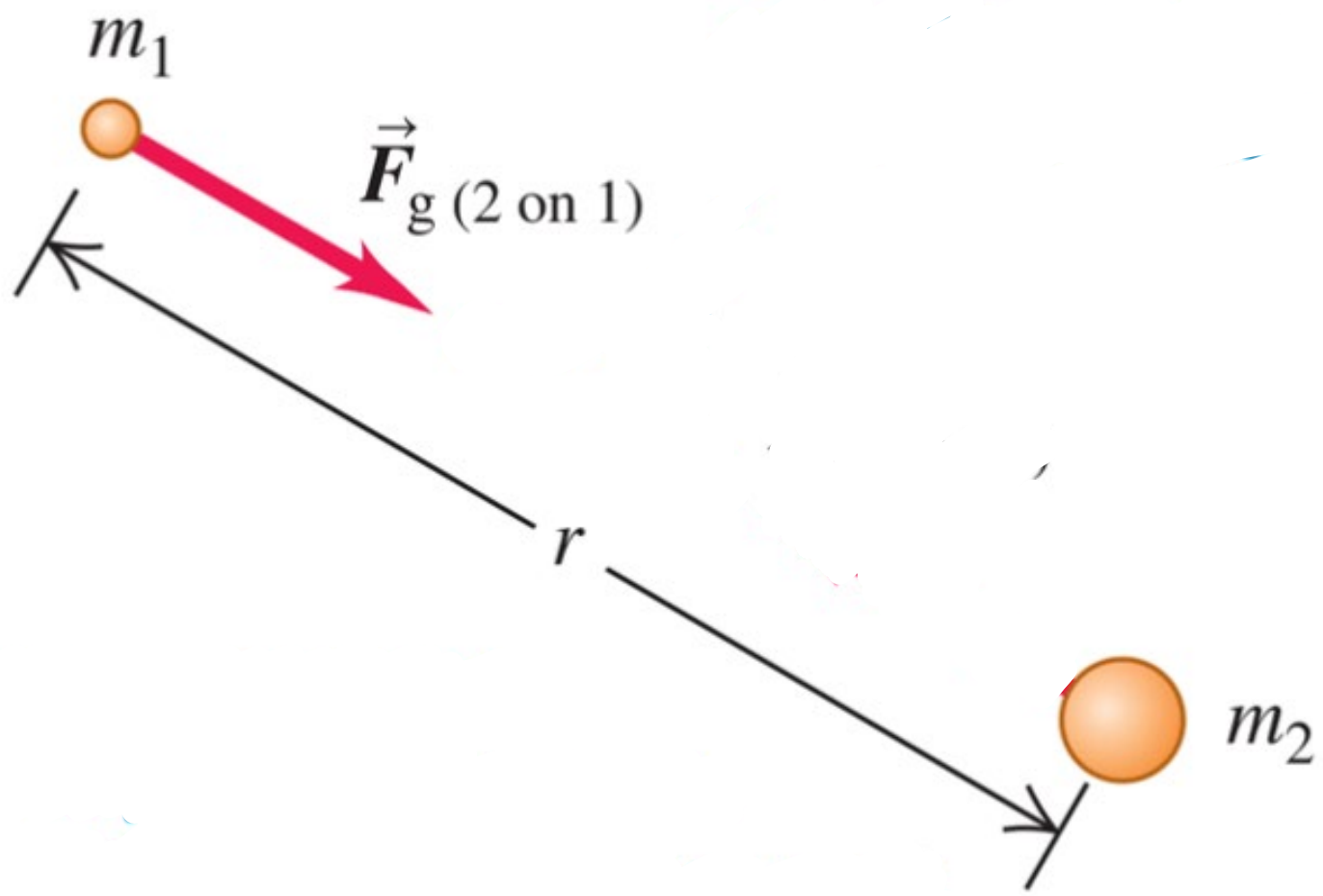
Distance between particles

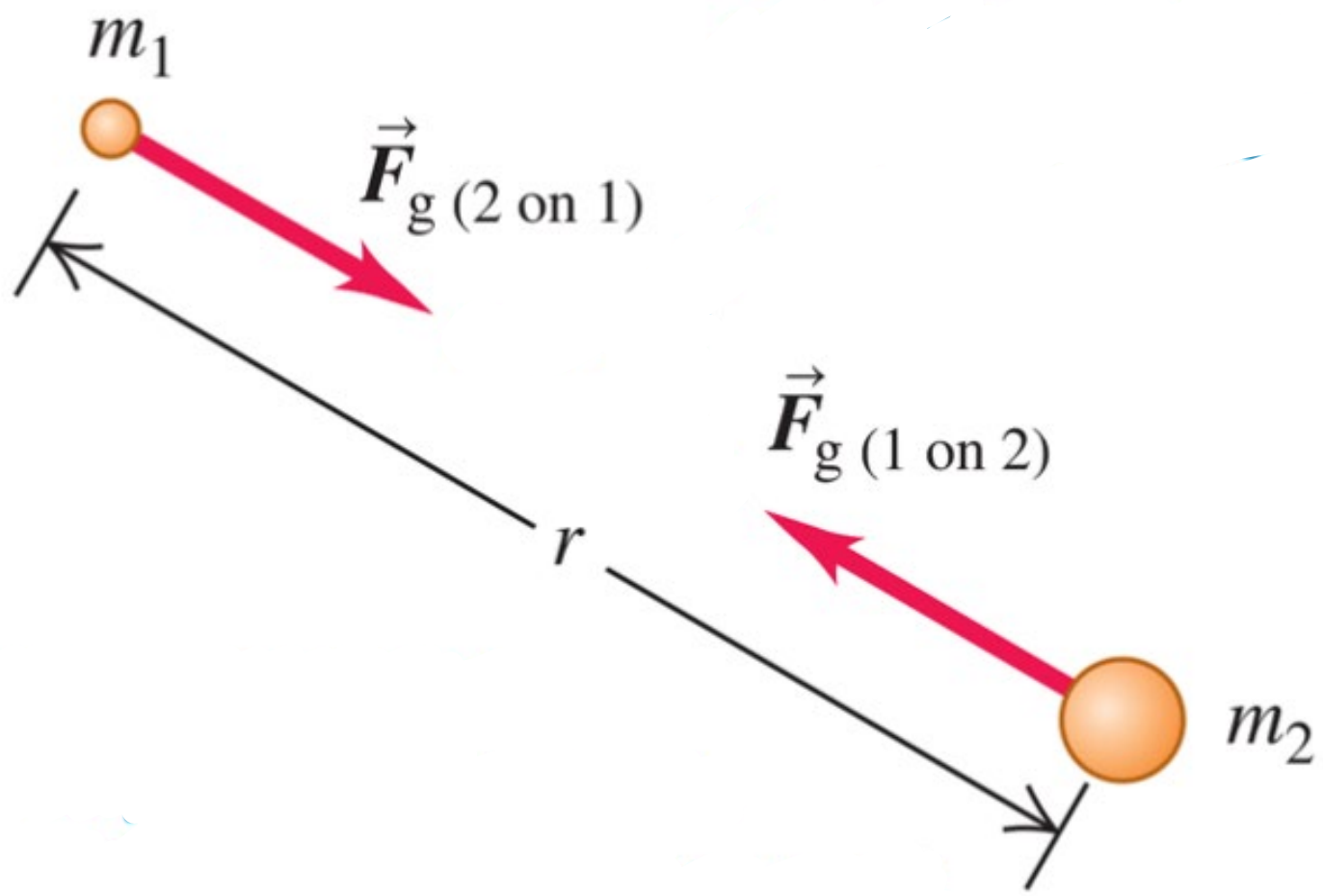
m_1



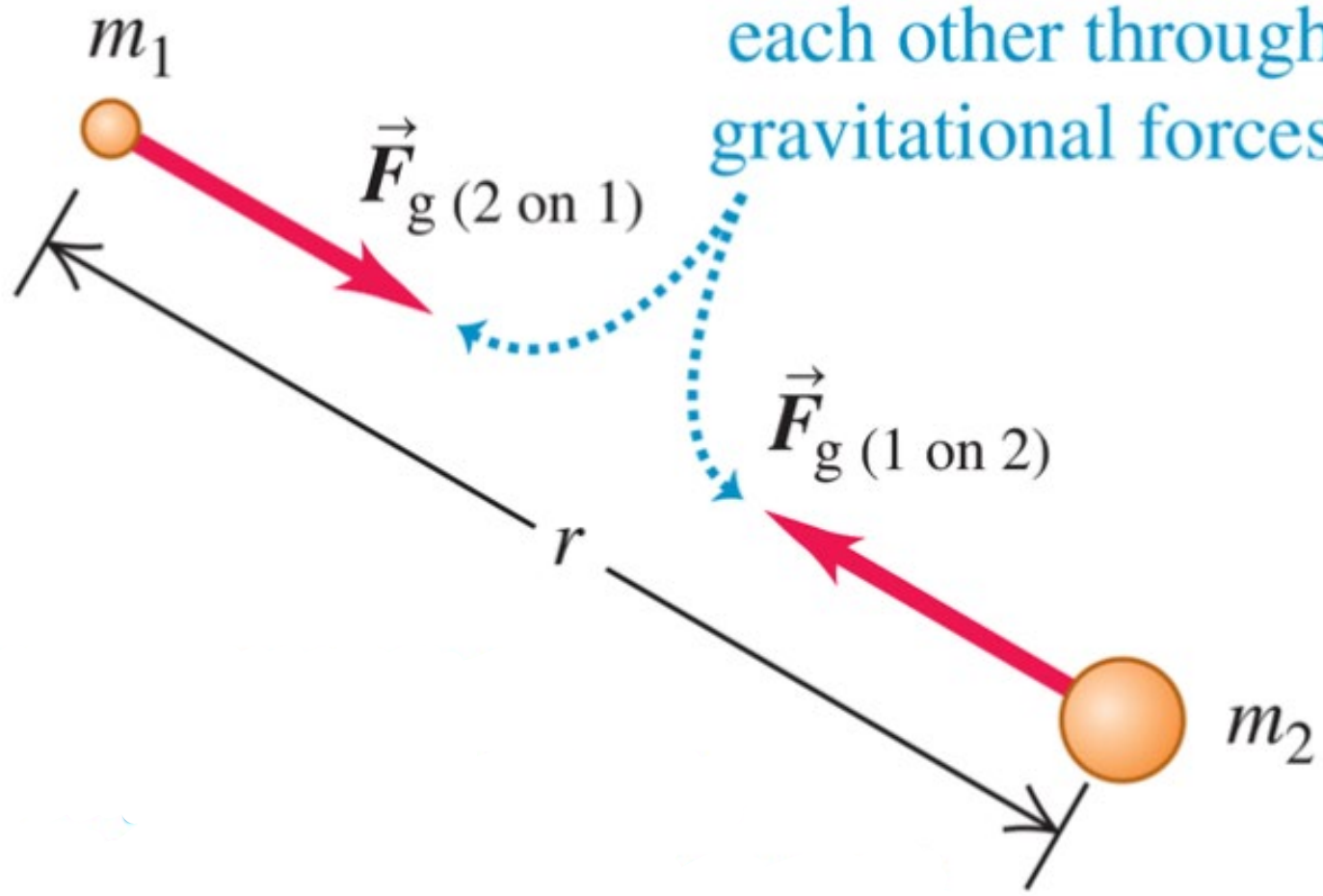
m_2



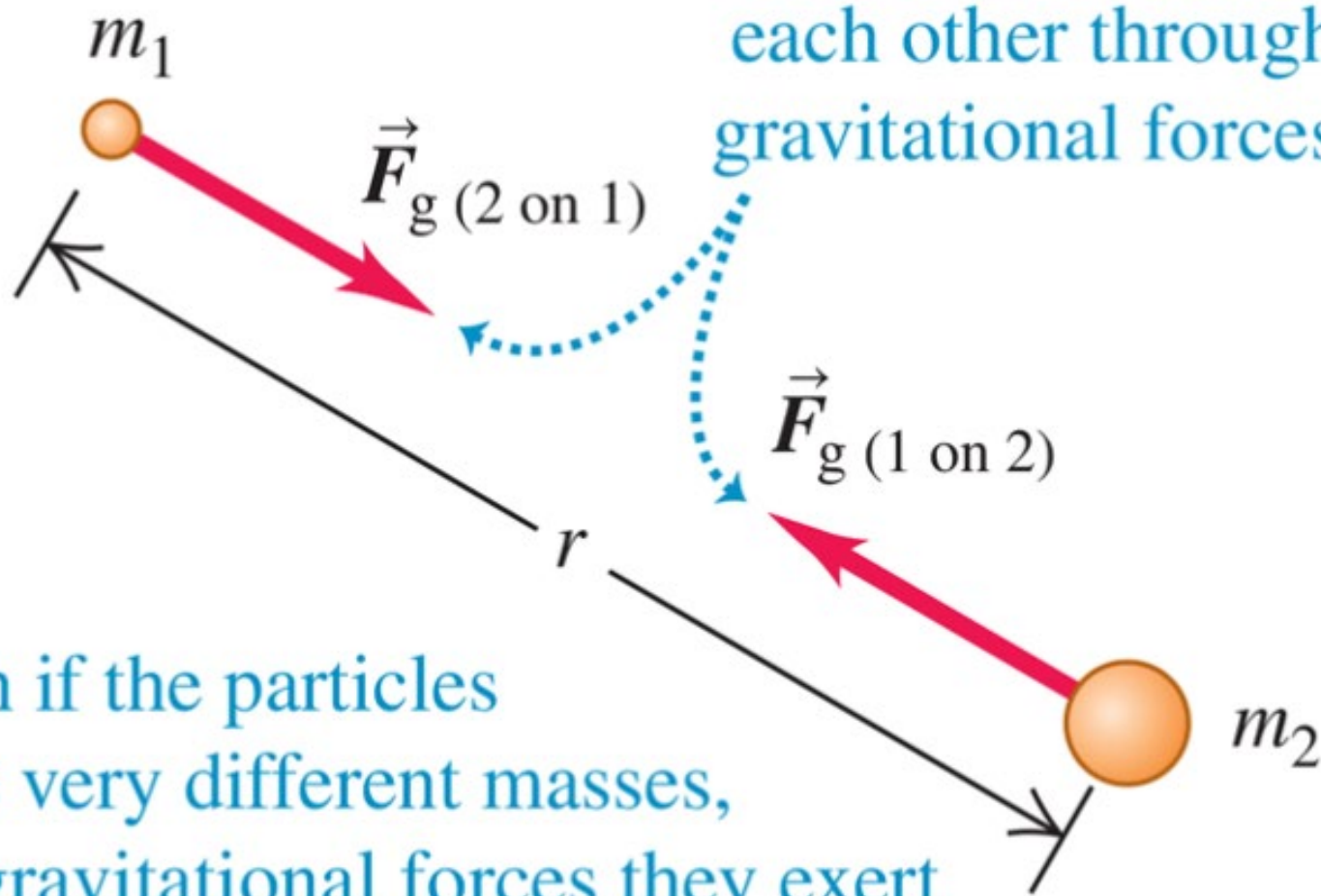




Any two particles attract each other through gravitational forces.



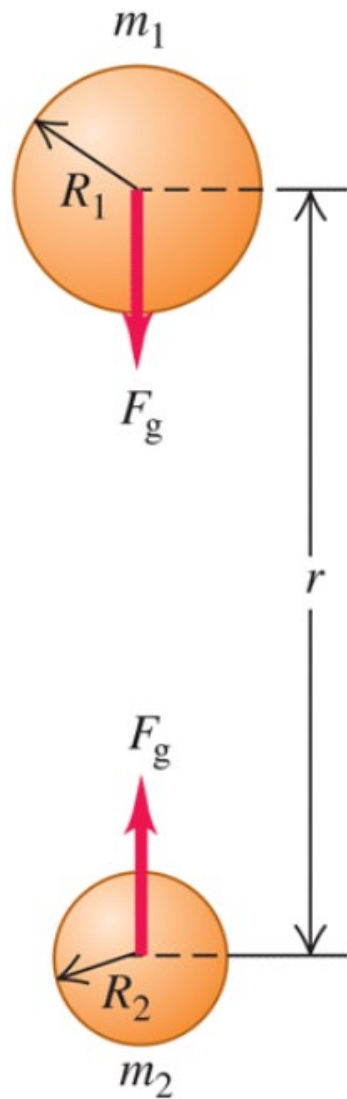
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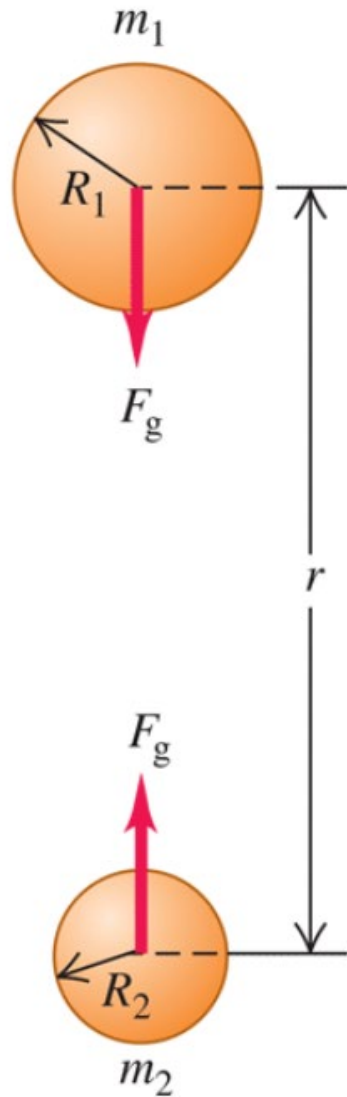
Even if the particles have very different masses, the gravitational forces they exert on each other are equal in magnitude:

$$F_{g(1 \text{ on } 2)} = F_{g(2 \text{ on } 1)}$$

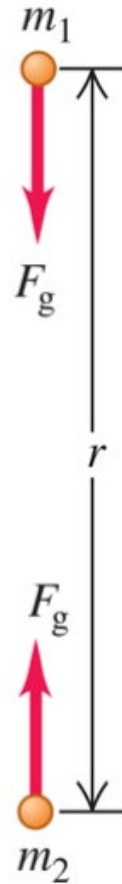
(a) The gravitational force between two spherically symmetric masses m_1 and m_2 ...

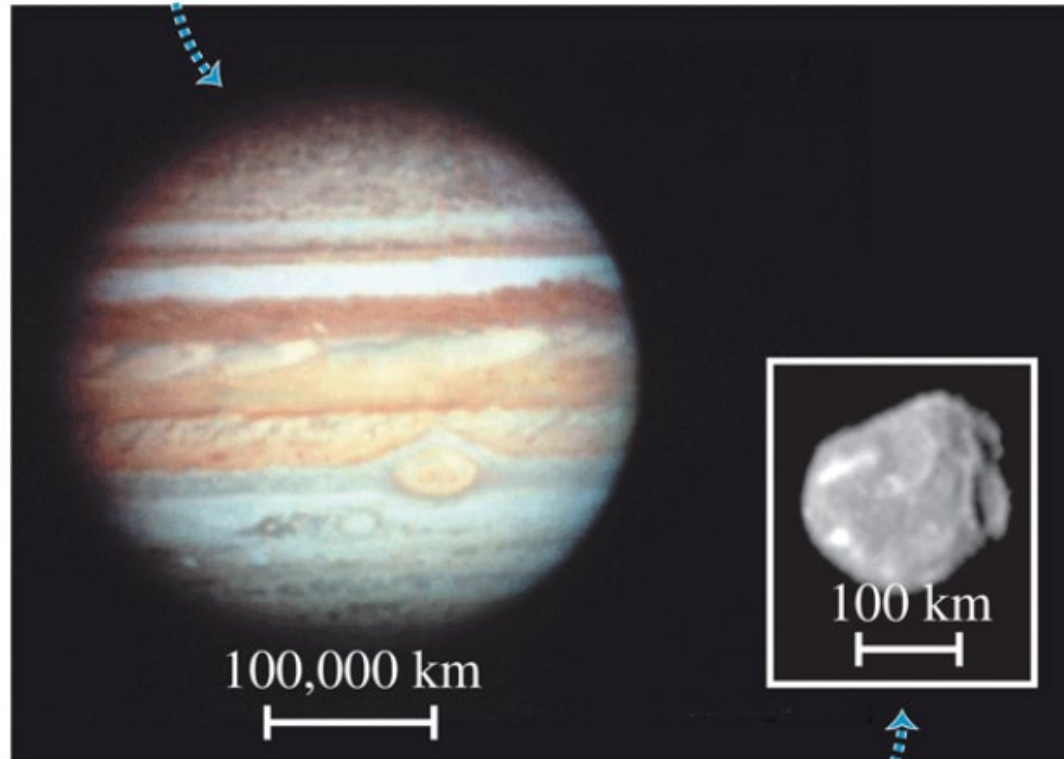


(a) The gravitational force between two spherically symmetric masses m_1 and m_2 ...

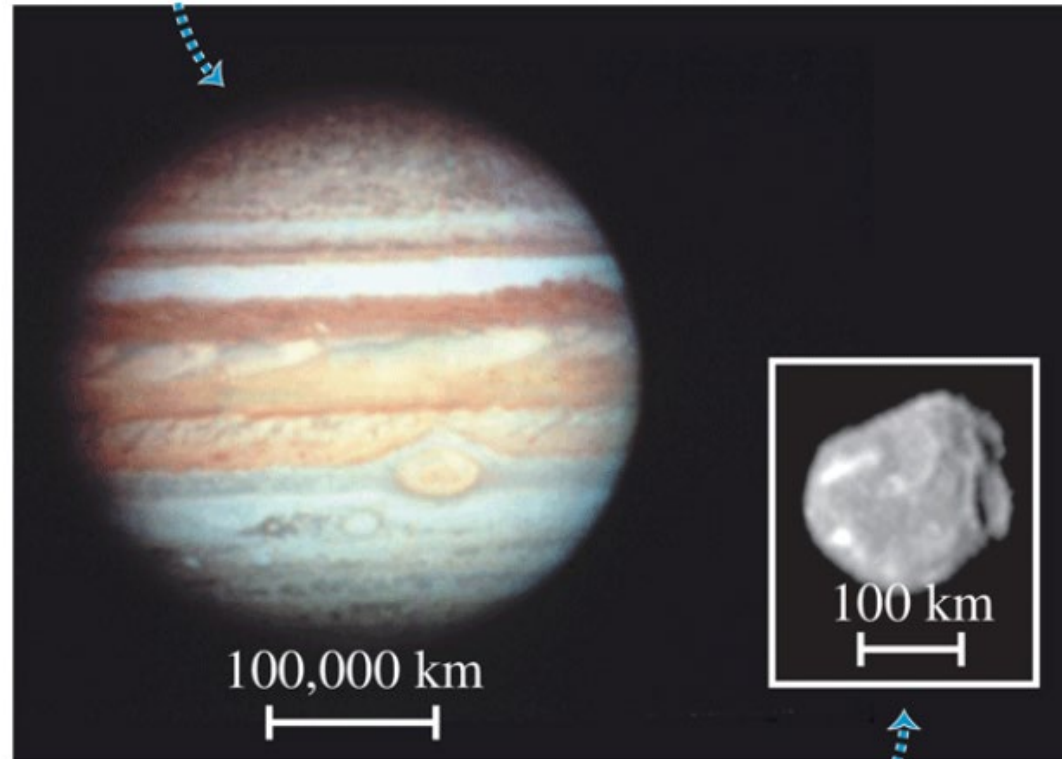


(b) ... is the same as if we concentrated all the mass of each sphere at the sphere's center.

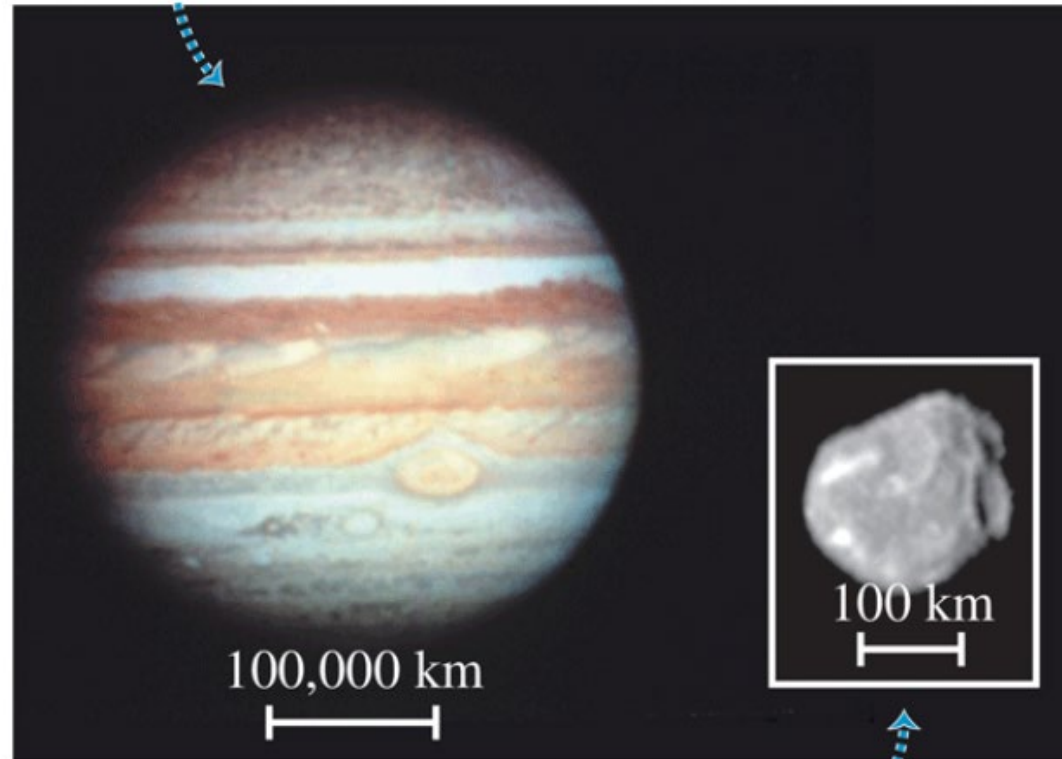




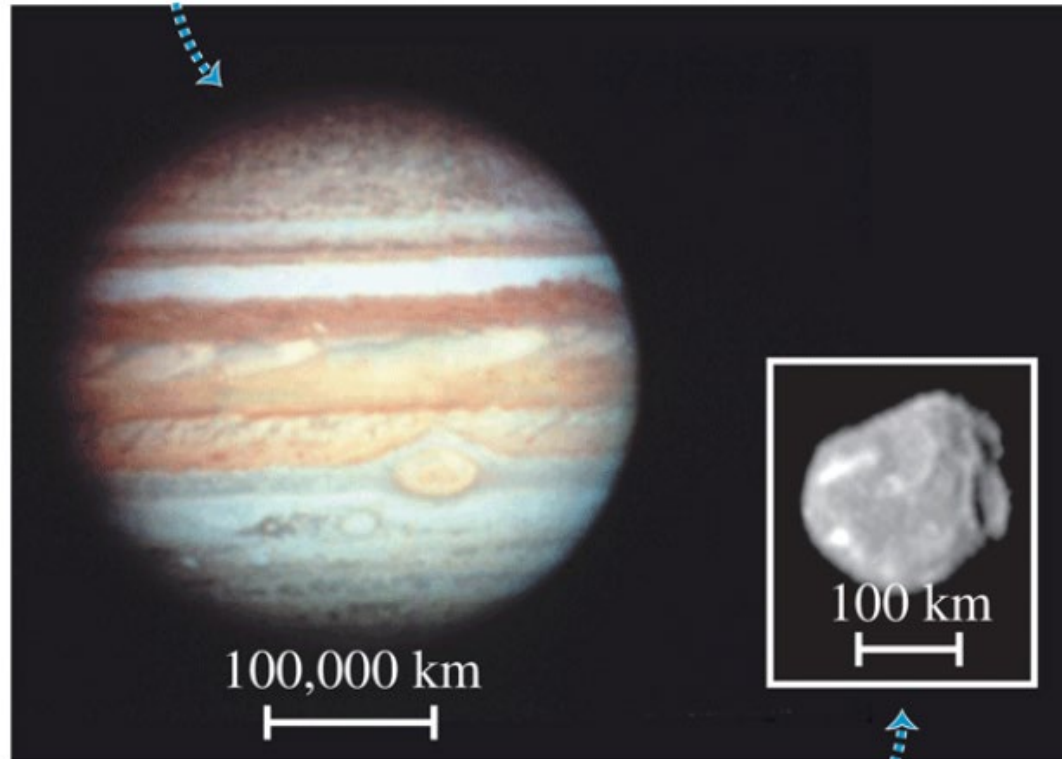
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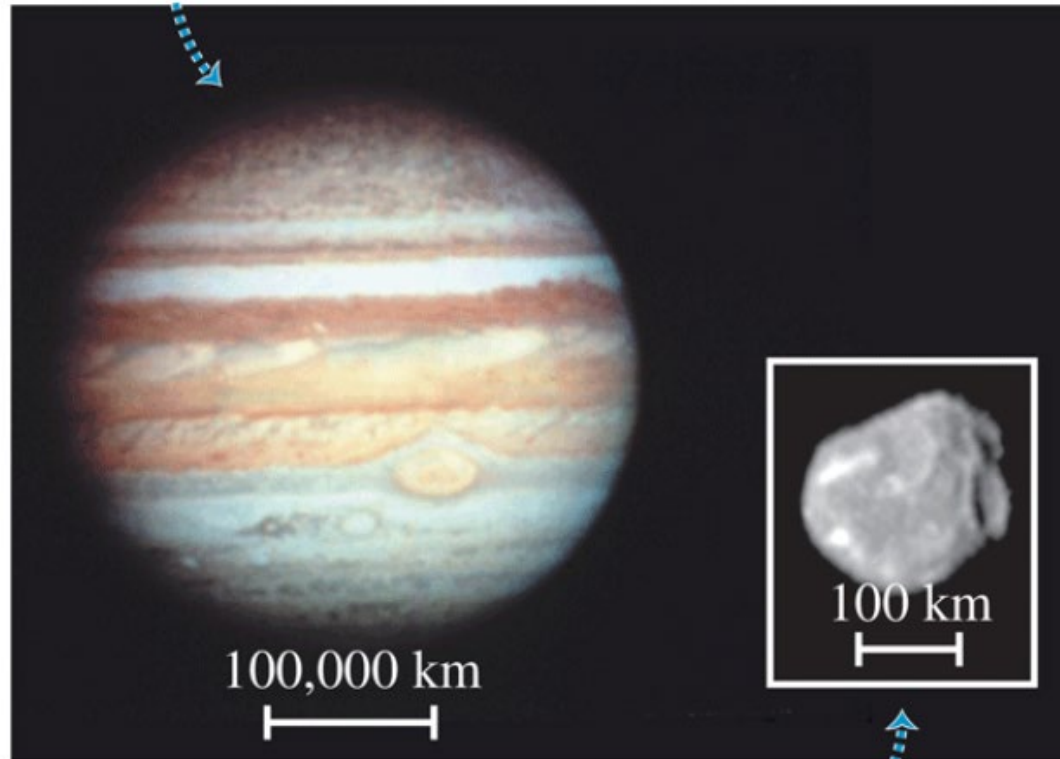


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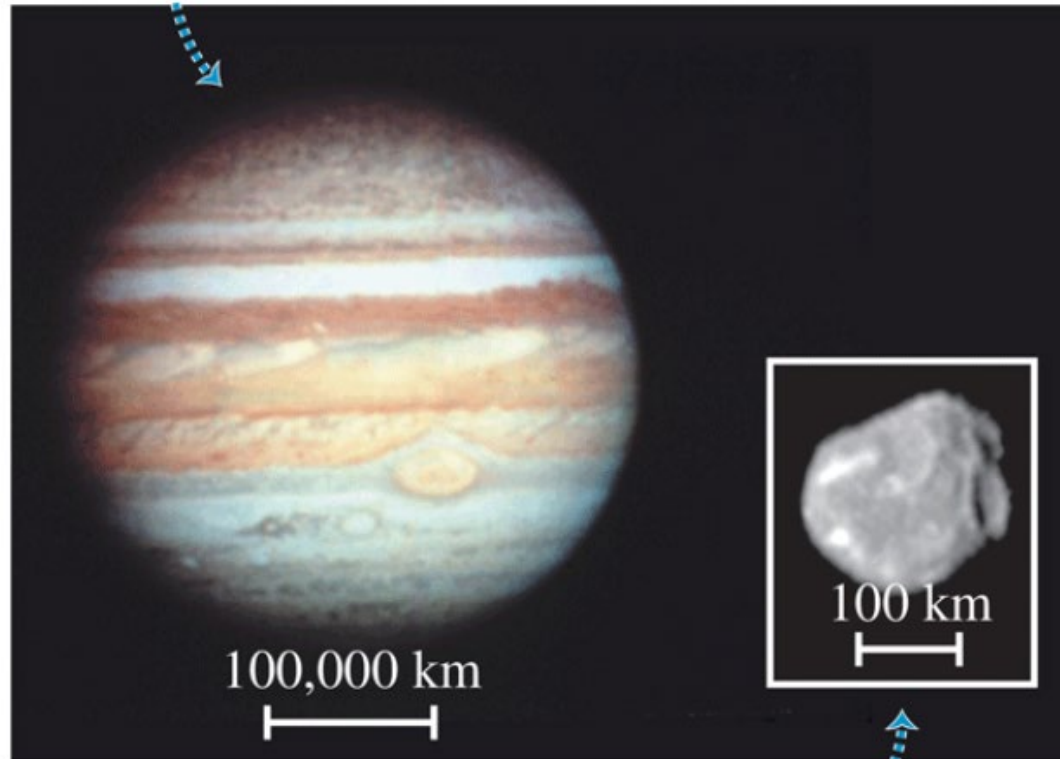
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Amalthea, one of Jupiter's small moons, has a relatively tiny mass (7.17×10^{18} kg, only about 3.8×10^{-9} the mass of Jupiter) and weak mutual gravitation, so it has an irregular shape.

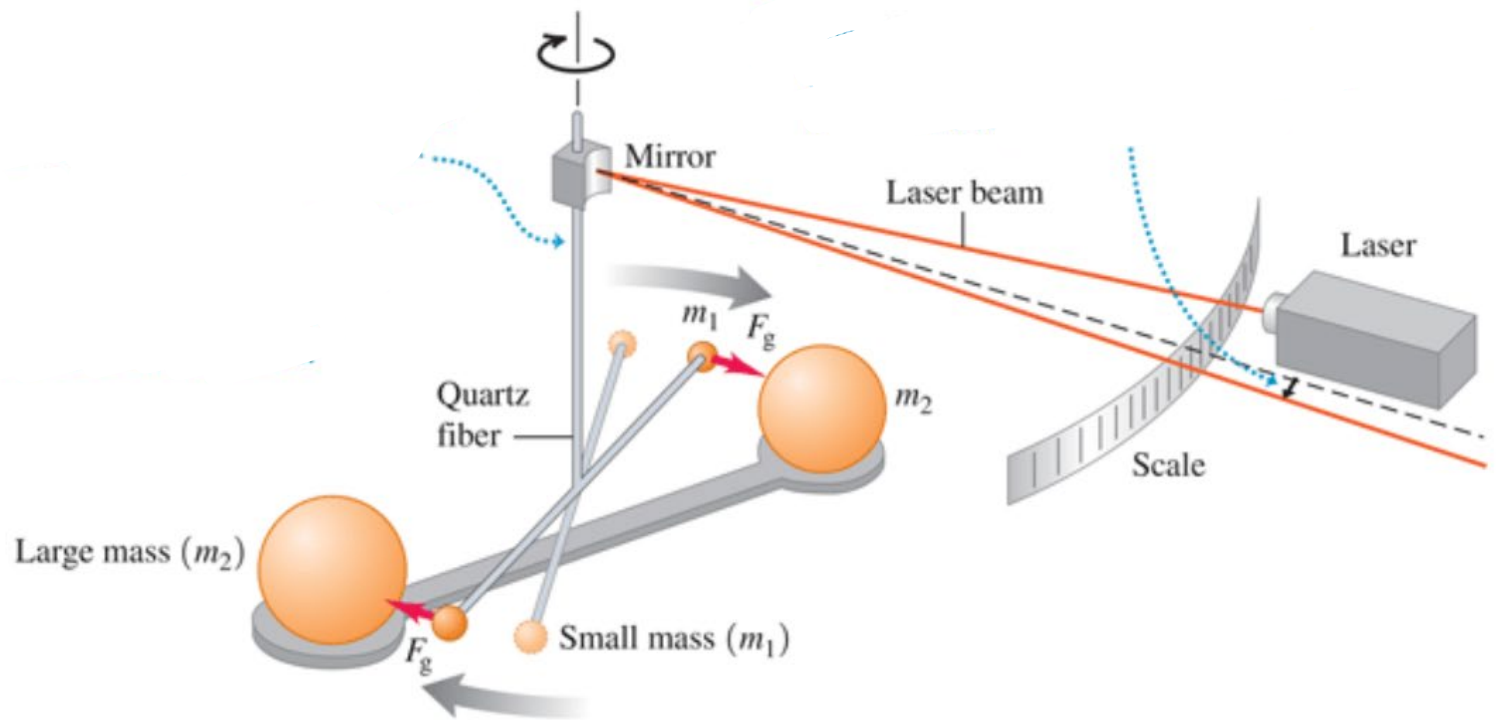
Determining the value of G

Determining the value of G

Cavendish torsion balance

Determining the value of G

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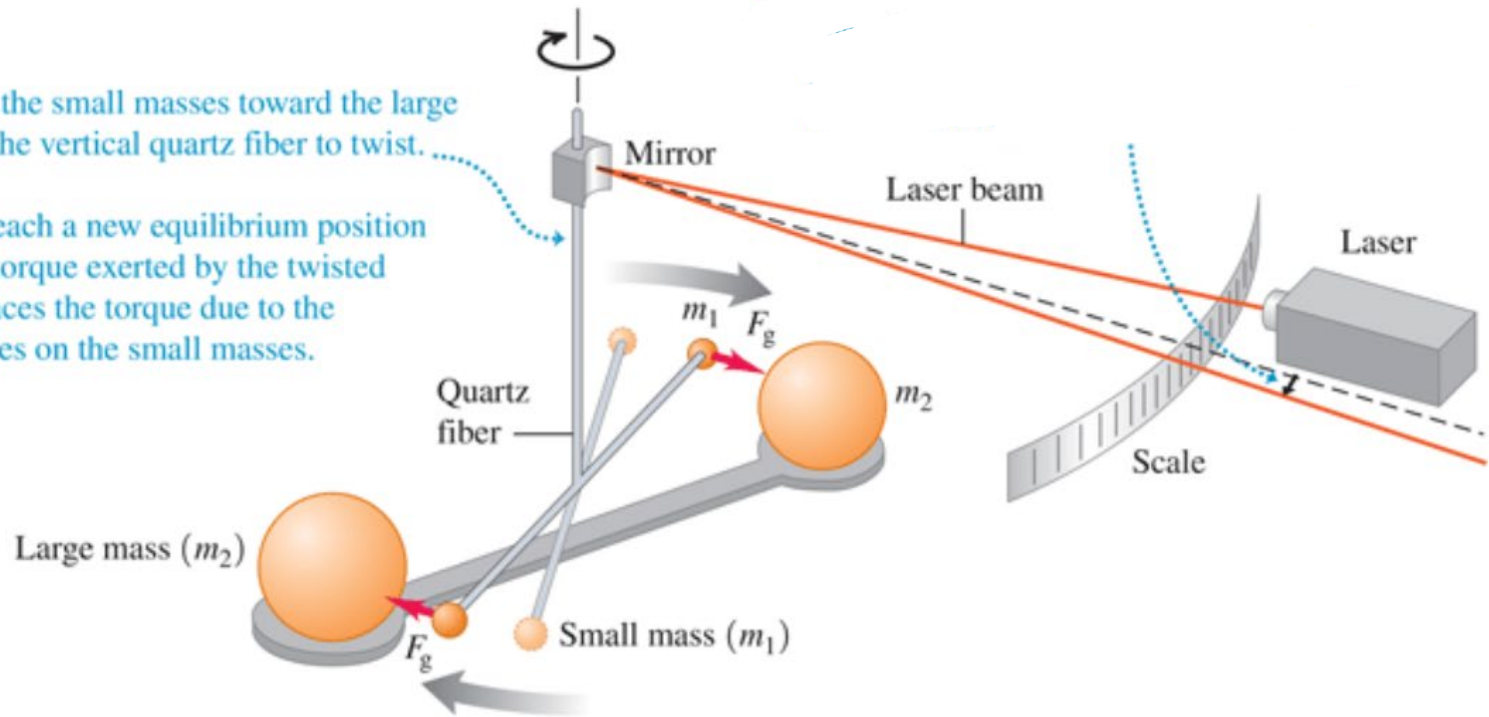


Determining the value of G

Cavendish torsion balance

- ① Gravitation pulls the small masses toward the large masses, causing the vertical quartz fiber to twist.

The small balls reach a new equilibrium position when the elastic torque exerted by the twisted quartz fiber balances the torque due to the gravitational forces on the small masses.

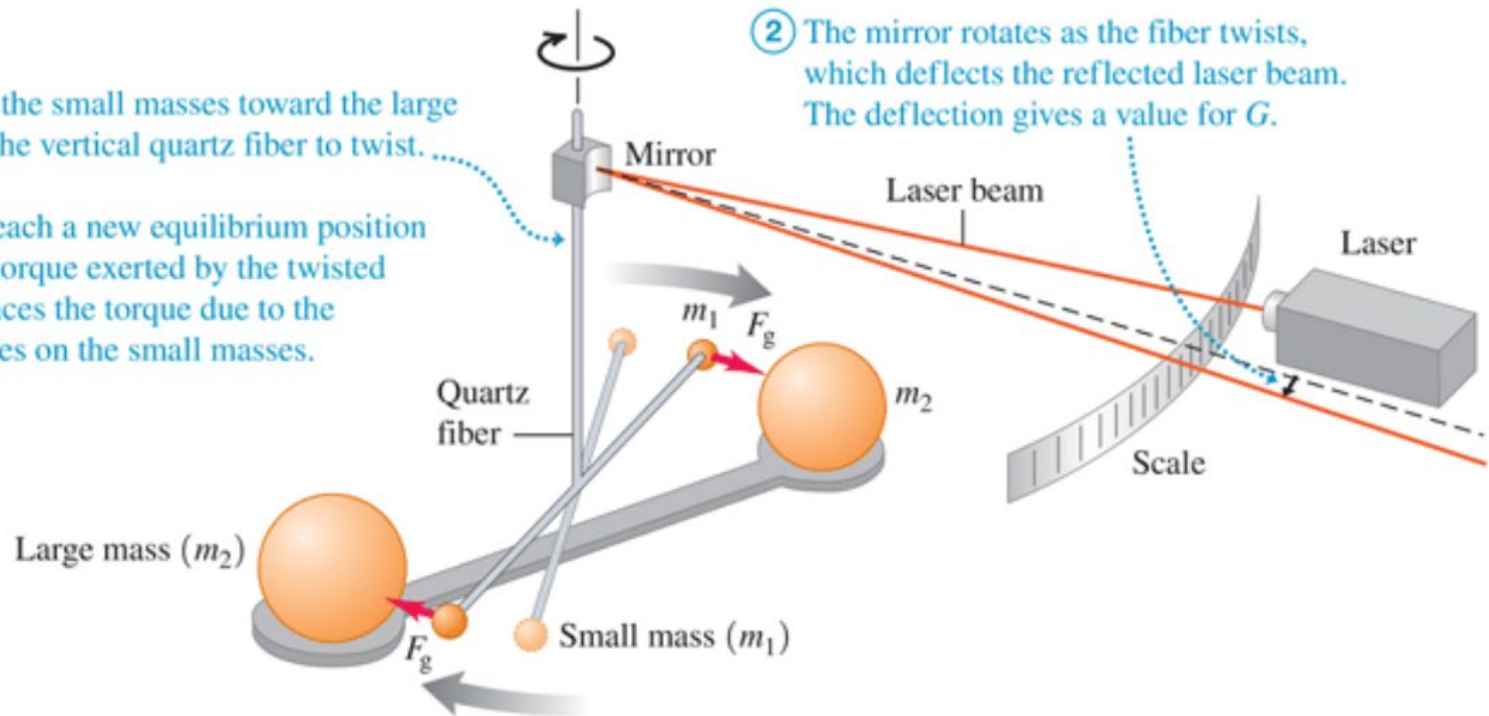


Determining the value of G

Cavendish torsion balance

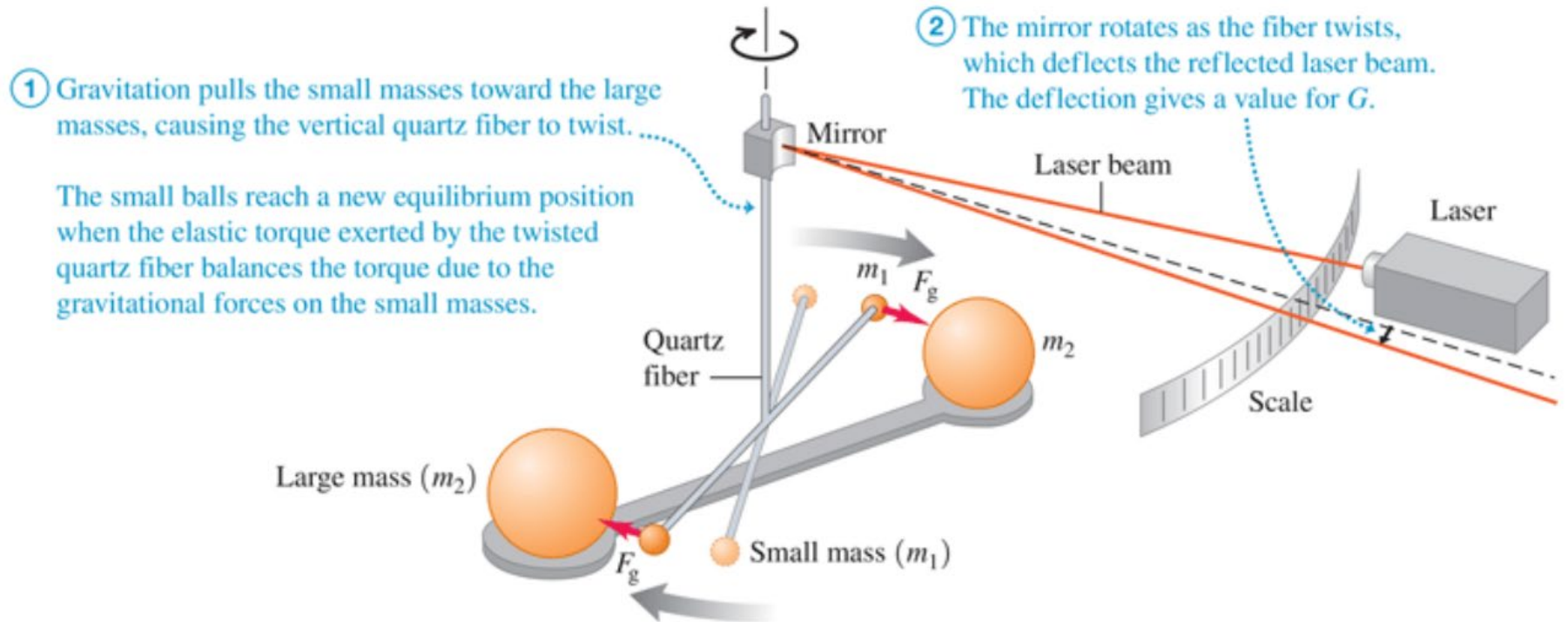
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Determining the value of G

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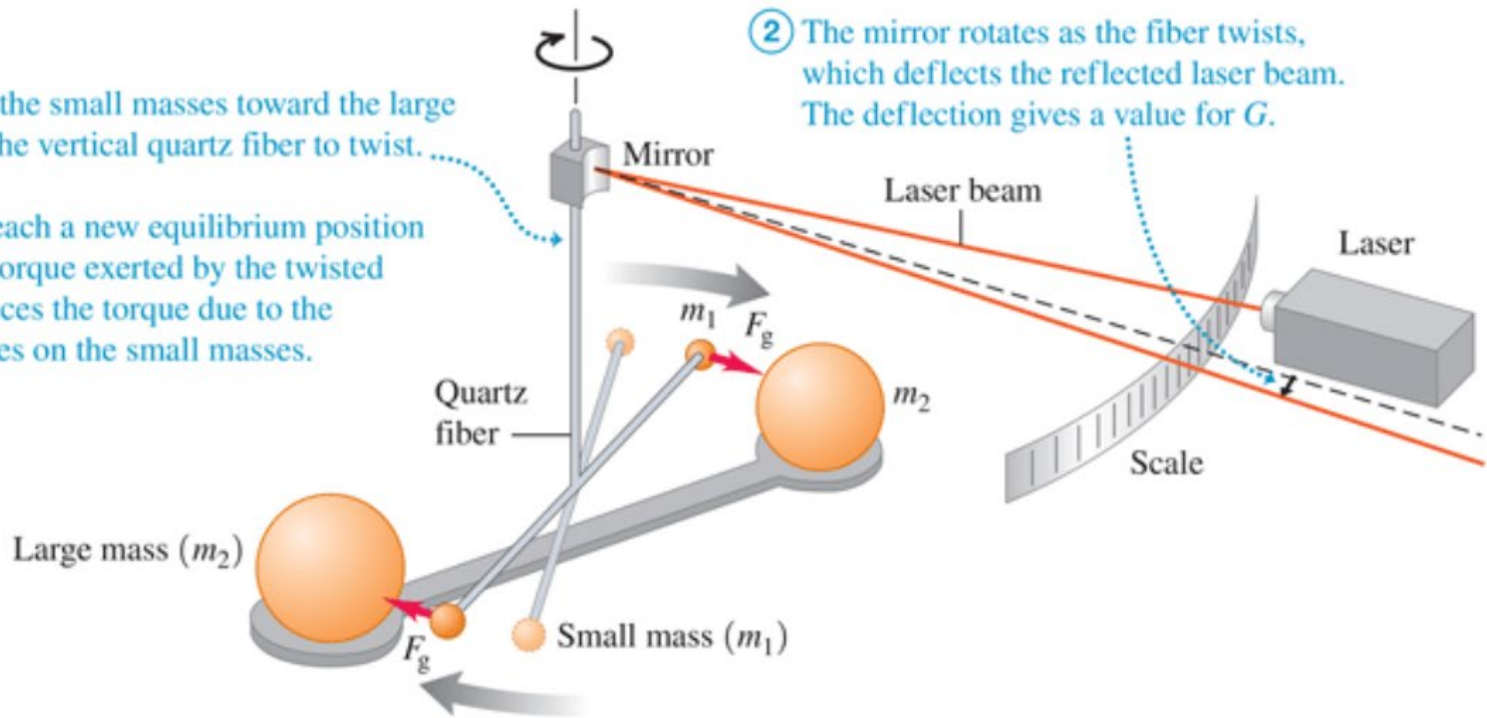
The principle of the Cavendish balance, used for determining the value of G . The angle of deflection has been exaggerated here for clarity.

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The principle of the Cavendish balance, used for determining the value of G. The angle of deflection has been exaggerated here for clarity.



$$G = 6.67408(31) * 10^{-11} \frac{N \cdot m^2}{kg^2}$$

The mass m_1 of one of the small spheres of a Cavendish balance is 0.0100 kg, the mass m_2 of the nearest large sphere is 0.500 kg, and the center-to-center distance between them is 0.0500 m. Find the gravitational force F_g on each sphere due to the other.

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$$F_g = \frac{(6.67 \times 10^{-11})(1 \times 10^{-2})(5 \times 10^{-1})}{25 \times 10^{-4}} \text{ N}$$

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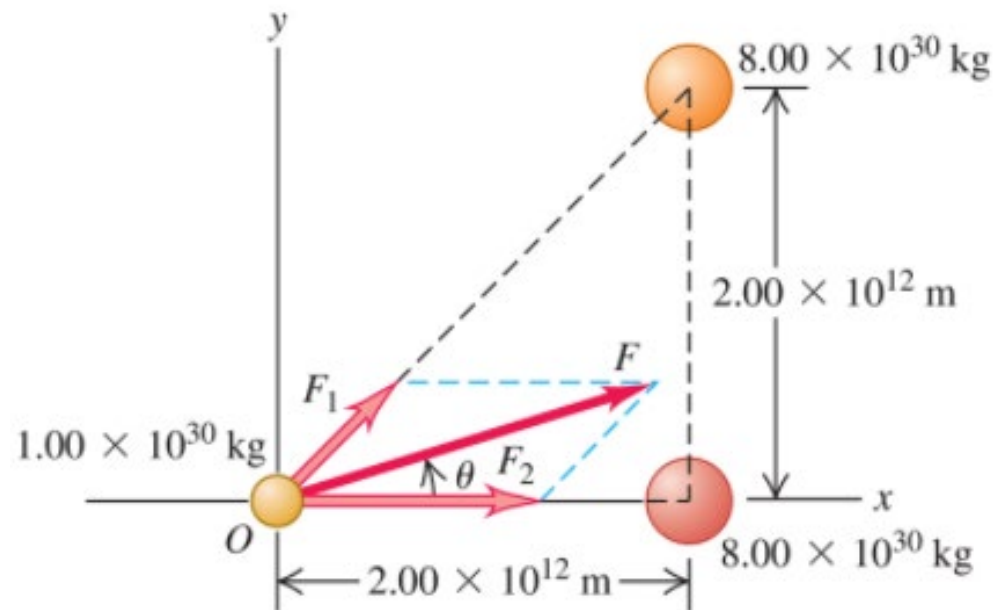
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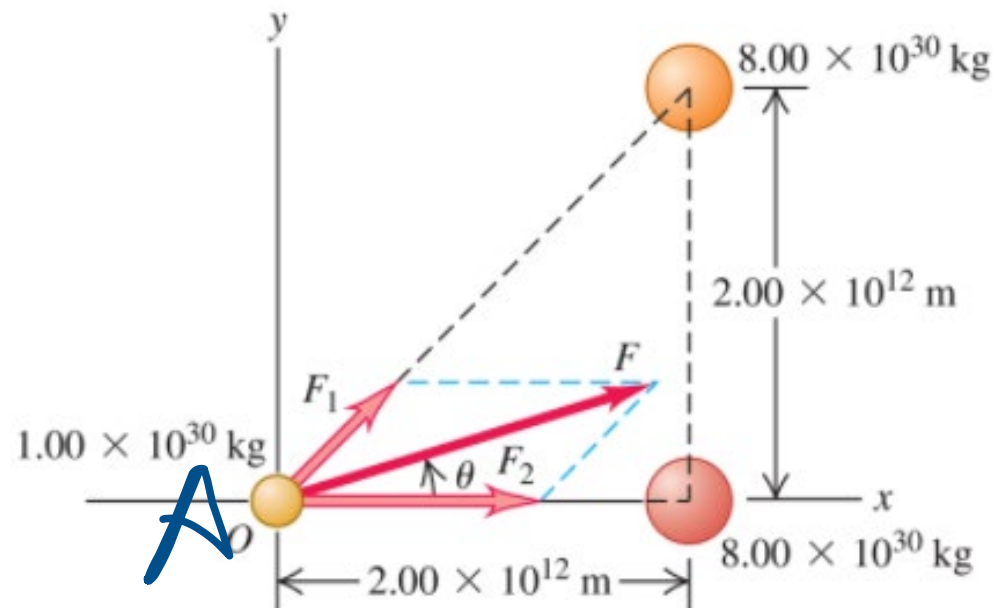
$$\& \quad a_1 = \frac{F_{2 \rightarrow 1}}{m_1} = \frac{1.33 \times 10^{-10} \text{ N}}{1 \times 10^{-2} \text{ kg}}$$

$$\Rightarrow \quad a_1 = 1.33 \times 10^{-8} \frac{\text{m}}{\text{s}^2}$$

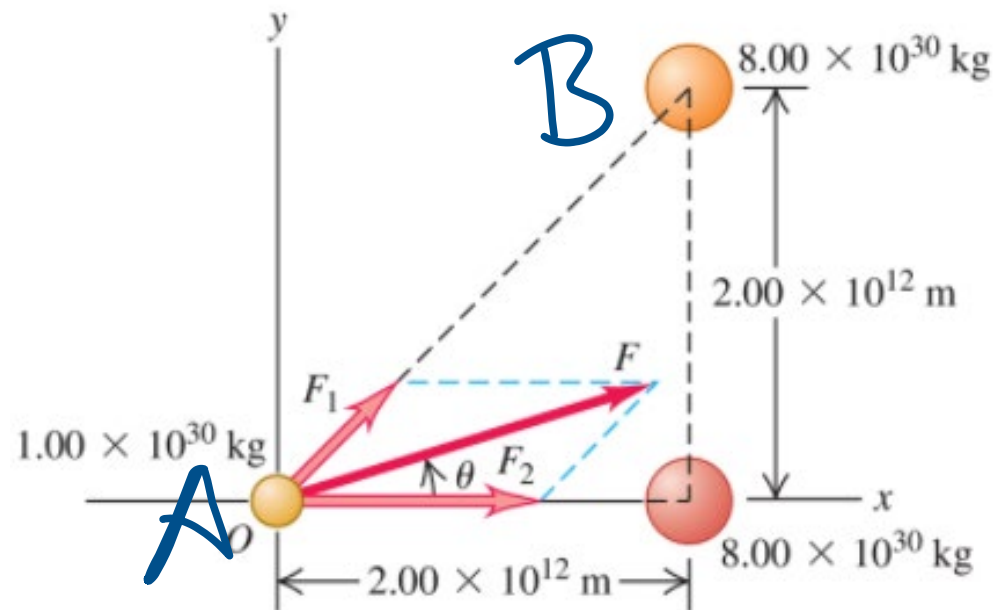
Many stars belong to *systems* of two or more stars held together by their mutual gravitational attraction. **Figure 13.5** shows a three-star system at an instant when the stars are at the vertices of a 45° right triangle. Find the total gravitational force exerted on the small star by the two large ones.



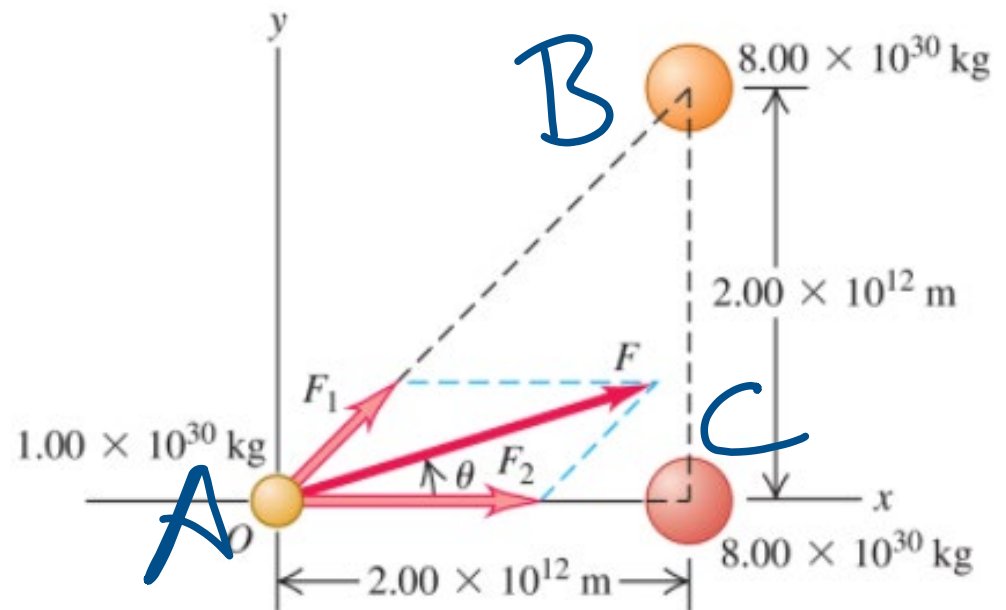
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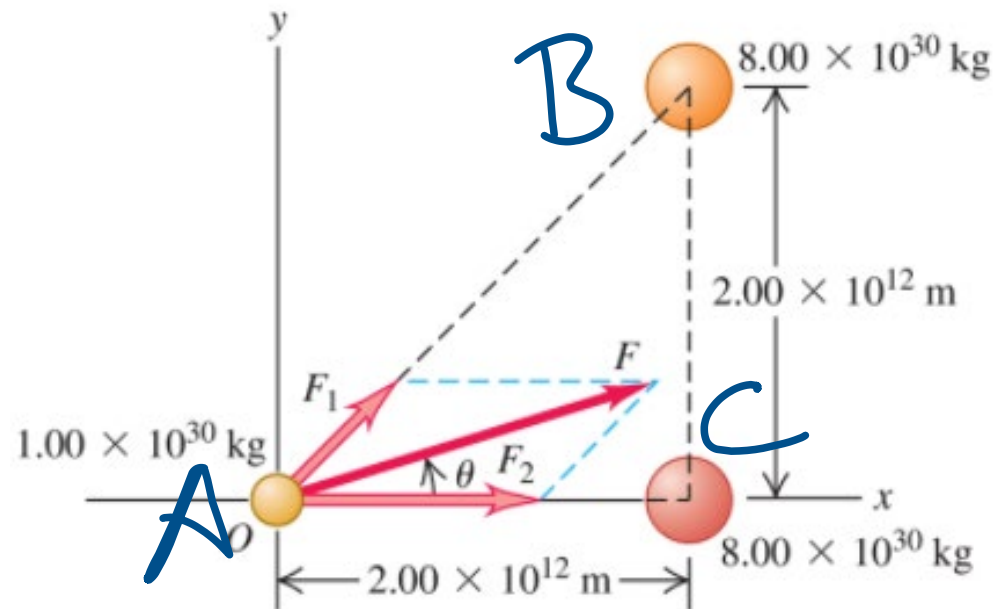


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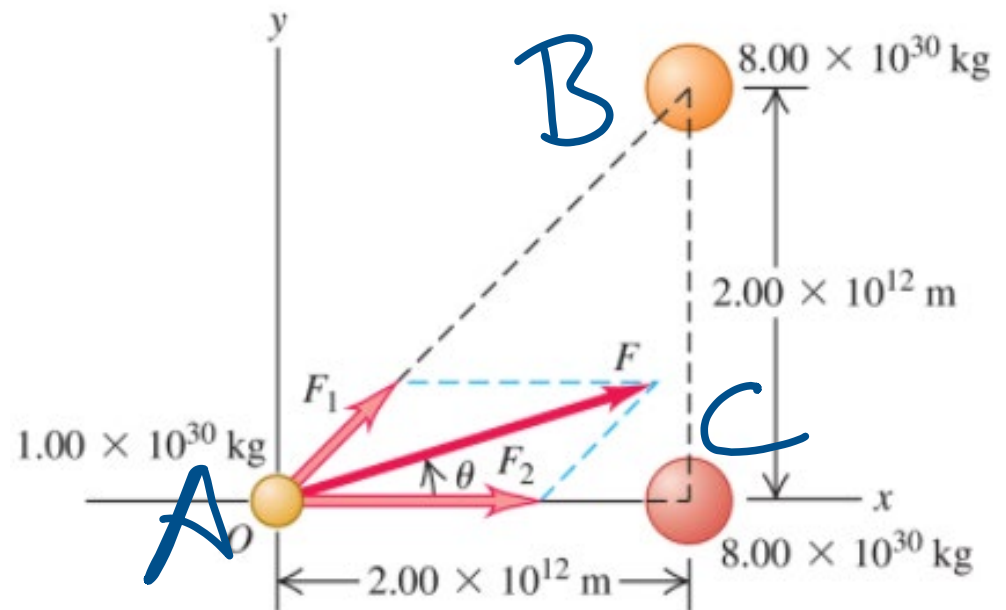
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$$\vec{F}_{B \rightarrow A} = \left(G \frac{M_B M_A}{Q_{AB}^2} \right) (\hat{i} \cos 45^\circ + \hat{j} \sin 45^\circ)$$

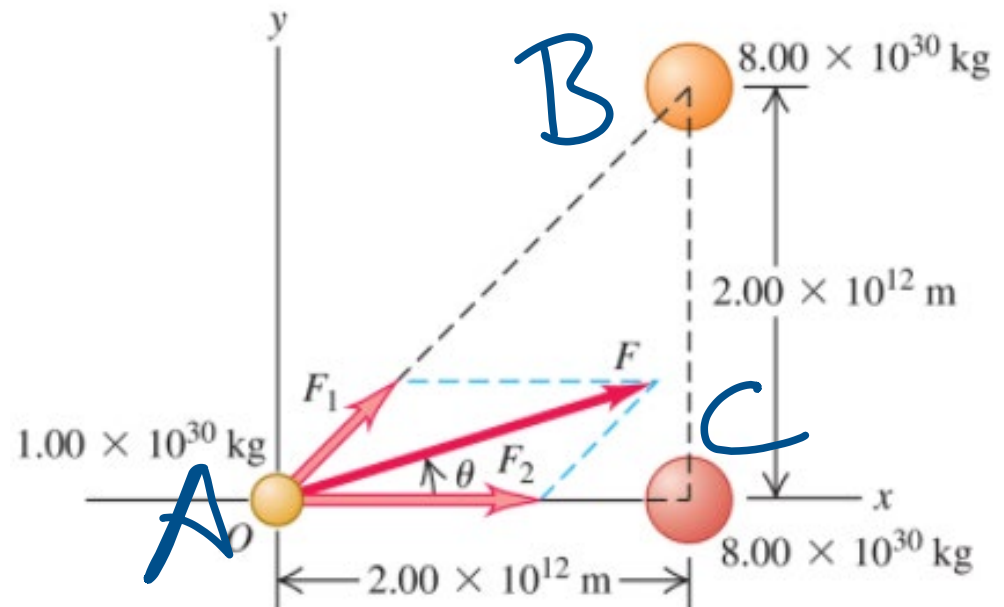


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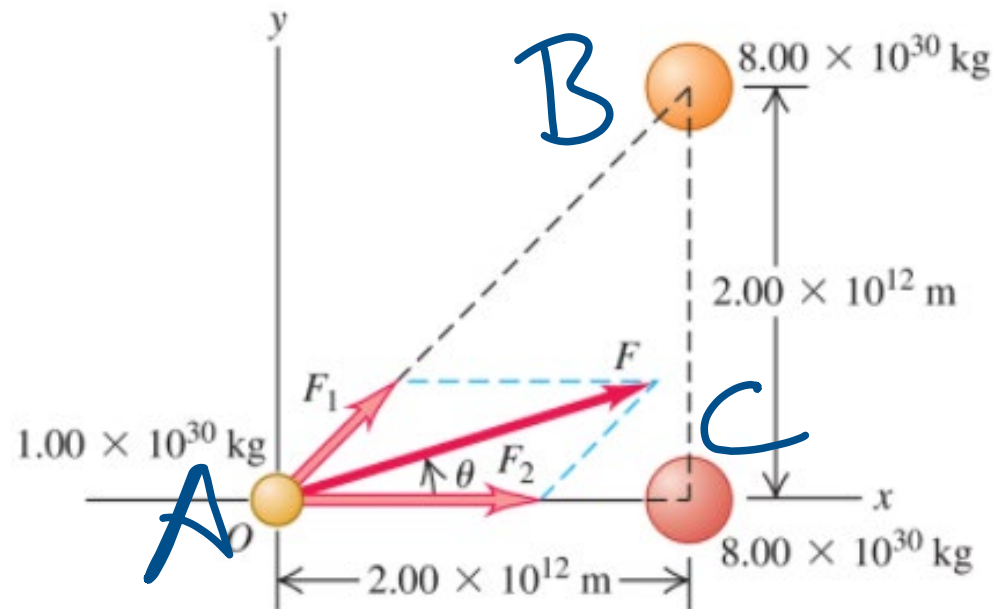
Many stars belong to *systems* of two or more stars held together by their mutual gravitational attraction. **Figure 13.5** shows a three-star system at an instant when the stars are at the vertices of a 45° right triangle. Find the total gravitational force exerted on the small star by the two large ones.

$$Q_{AB}^2 = Q_{AC}^2 + Q_{CB}^2$$

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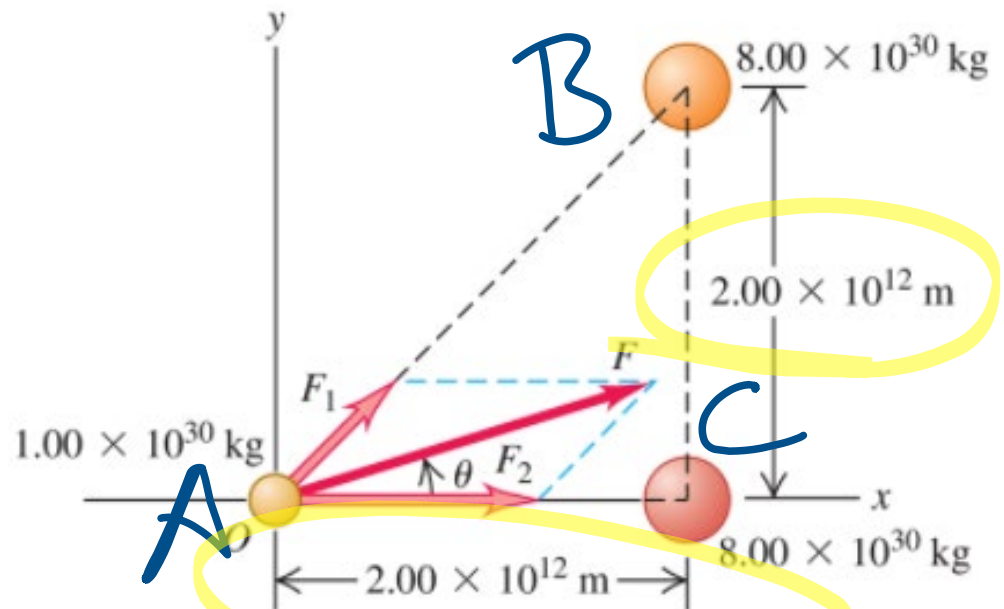
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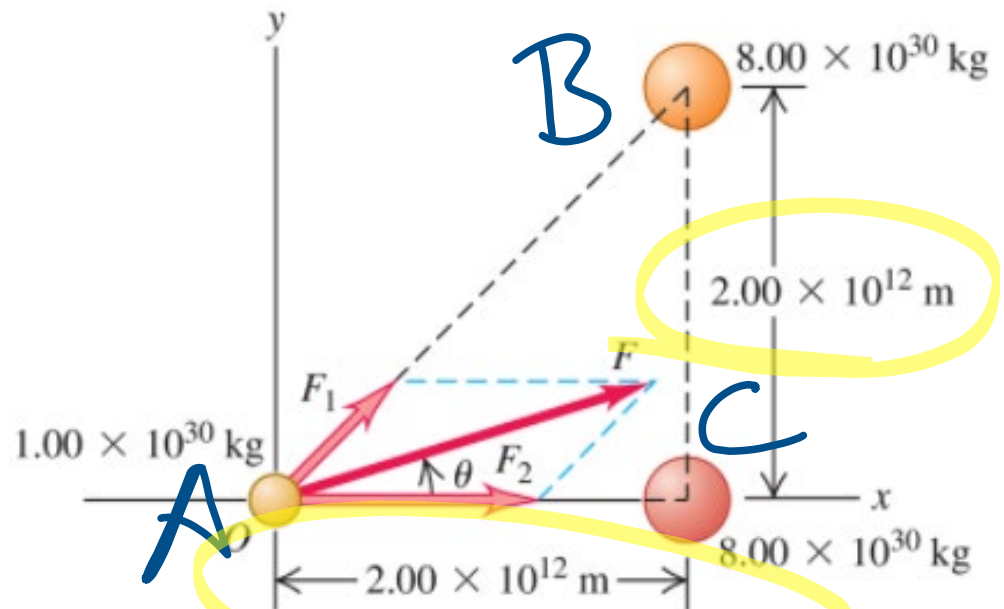
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$$\Rightarrow F_A = \sqrt{(F_A)_x^2 + (F_A)_y^2} = (8^2 + 4.7^2)^{\frac{1}{2}} 10^{25} \text{ N}$$

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$$\Rightarrow \Theta = 14.6^\circ$$

§ 13.2

§13.2

Weight

§ 13.2

Weight

Object weigh
about 17%
of that on
earth



$$w = F_g = \frac{Gm_E m}{R_E^2}$$

Weight of an object
at the earth's surface ...

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Weight of an object
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$$w = F_g = \frac{Gm_E m}{R_E^2}$$

... equals gravitational force
the earth exerts on object.

Gravitational constant

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Gravitational constant \cdots

Weight of an object at the earth's surface ... $w = F_g = \frac{Gm_E m}{R_E^2}$ \cdots Mass of the earth

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Weight of an object at the earth's surface ...
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$$w = F_g = \frac{Gm_E m}{R_E^2}$$

Gravitational constant
Mass of the earth
Mass of object

Weight of an object at the earth's surface ...
... equals gravitational force the earth exerts on object.

$$w = F_g = \frac{Gm_E m}{R_E^2}$$

Gravitational constant
Mass of the earth
Mass of object
Radius of the earth

Weight of an object at the earth's surface ...
... equals gravitational force the earth exerts on object.

$$w = F_g = \frac{Gm_E m}{R_E^2}$$

Gravitational constant
Mass of the earth
Mass of object
Radius of the earth

$$g = \frac{Gm_E}{R_E^2}$$

Weight of an object at the earth's surface ...
 ... equals gravitational force the earth exerts on object.

$$w = F_g = \frac{Gm_E m}{R_E^2}$$

Gravitational constant $\rightarrow G$
 Mass of the earth $\rightarrow m_E$
 Mass of object $\rightarrow m$
 Radius of the earth $\rightarrow R_E$

Acceleration due to gravity at the earth's surface

$$g = \frac{Gm_E}{R_E^2}$$

Weight of an object at the earth's surface ...
 ... equals gravitational force the earth exerts on object.

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Gravitational constant
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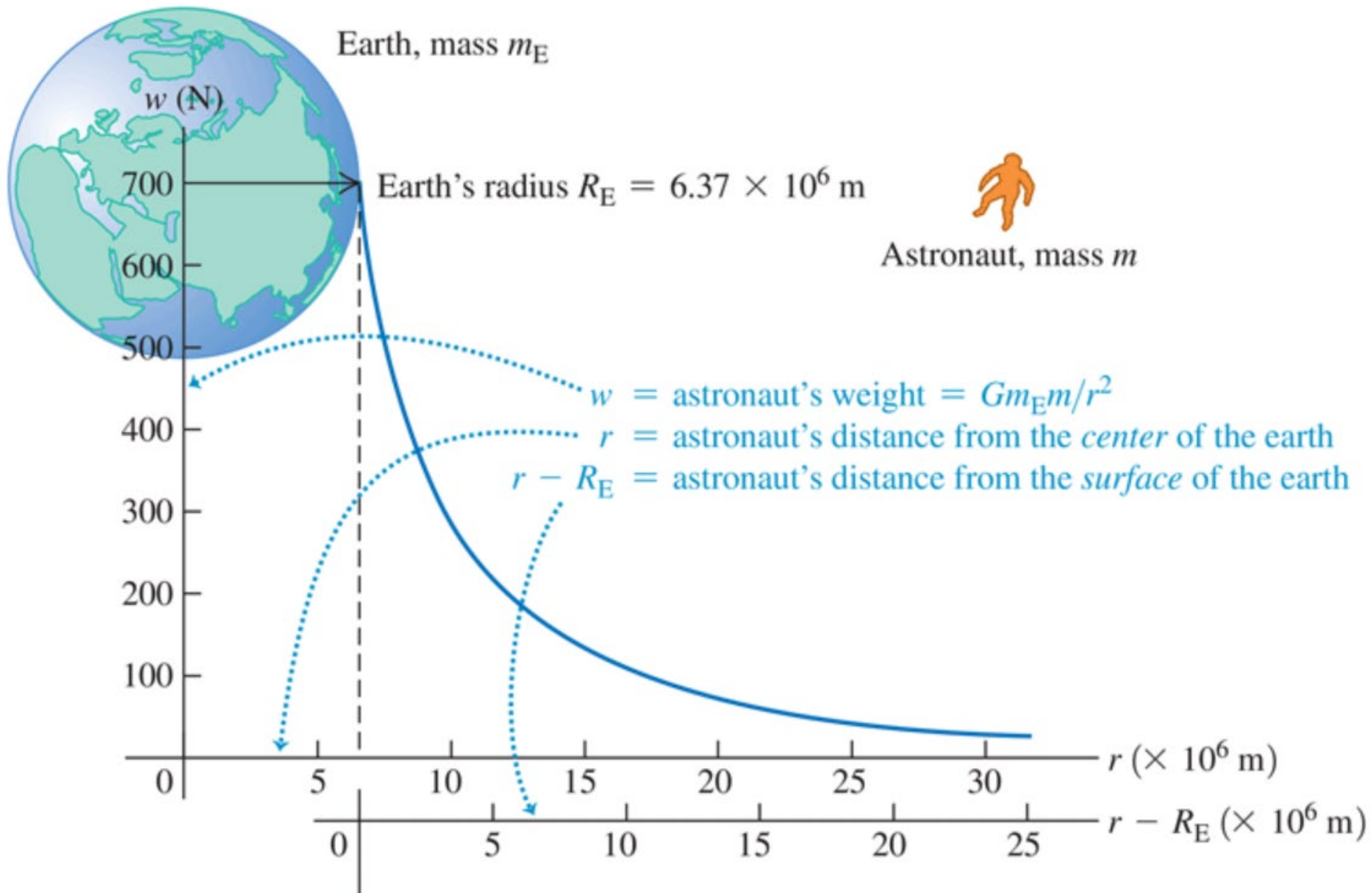
Acceleration due to gravity at the earth's surface

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Gravitational constant
Mass of the earth
Radius of the earth



In an airliner at high altitude, you are farther from the center of the earth than when on the ground and hence weigh slightly less. Can you show that at an altitude of 10 km above the surface, you weigh 0.3% less than you do on the ground?



Density

$$\rho = \frac{\text{mass}}{\text{Volume}}$$

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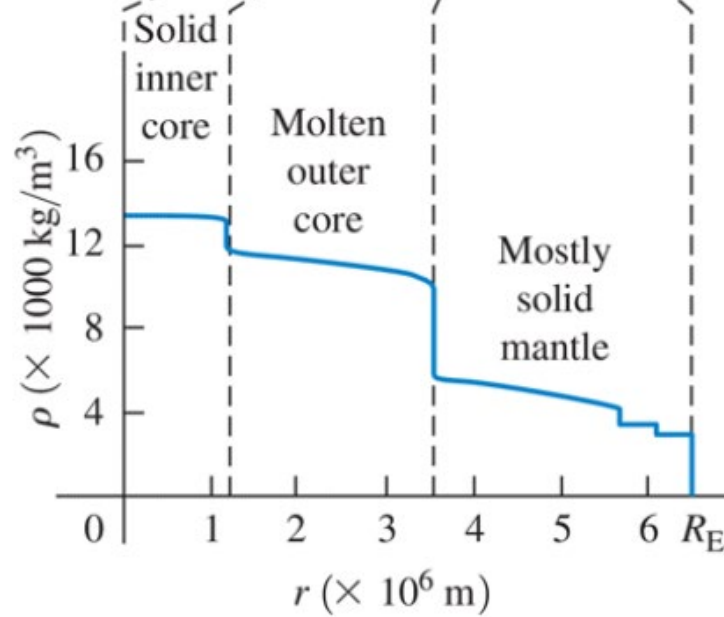
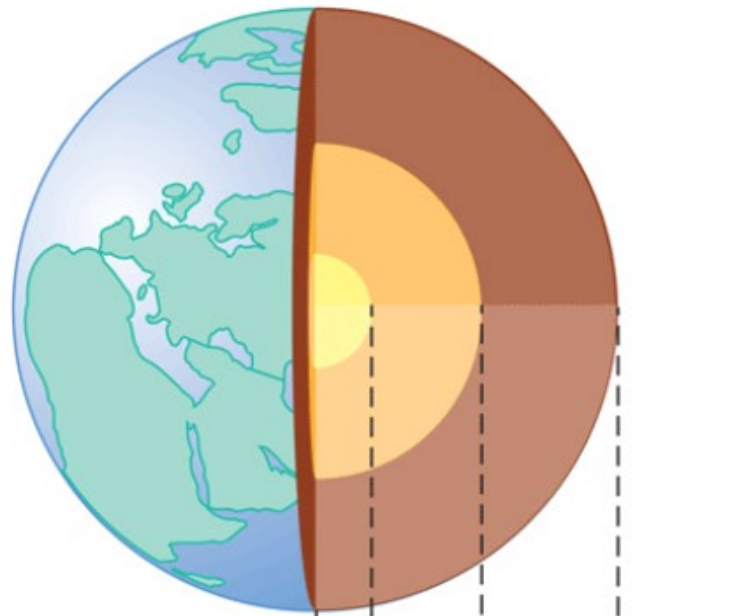
$$\rho_{\text{EARTH}} = 5500 \text{ kg/m}^3 = 5.5 \text{ g/cm}^3$$

Density

$$\rho = \frac{\text{mass}}{\text{Volume}}$$

$$\rho_{\text{EARTH}} = 5500 \text{ kg/m}^3 = 5.5 \text{ g/cm}^3$$

$$\rho_{\text{WATER}} = 1 \text{ g/cm}^3$$



A robotic lander with an earth weight of 3430 N is sent to Mars, which has radius $R_M = 3.39 \times 10^6$ m and mass $m_M = 6.42 \times 10^{23}$ kg (see [Appendix F](#)). Find the weight F_g of the lander on the Martian surface and the acceleration there due to gravity, g_M .



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$$\Rightarrow F_g = 1304 \times 10^{-11+23+12} \text{ N} = 1304 \text{ N} = 1.3 \times 10^3 \text{ N}$$

$$g_M = \frac{F_g}{m} = \left(\frac{1300}{3430/9.8}\right) \frac{\text{m}}{\text{s}^2}$$

A robotic lander with an earth weight of 3430 N is sent to Mars, which has radius

$R_M = 3.39 \times 10^6$ m and mass $m_M = 6.42 \times 10^{23}$ kg (see Appendix F). Find the weight F_g of the lander on the Martian surface and the acceleration there due to gravity, g_M .

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$$\Rightarrow g_M = 3.7 \frac{\text{m}}{\text{s}^2}$$

